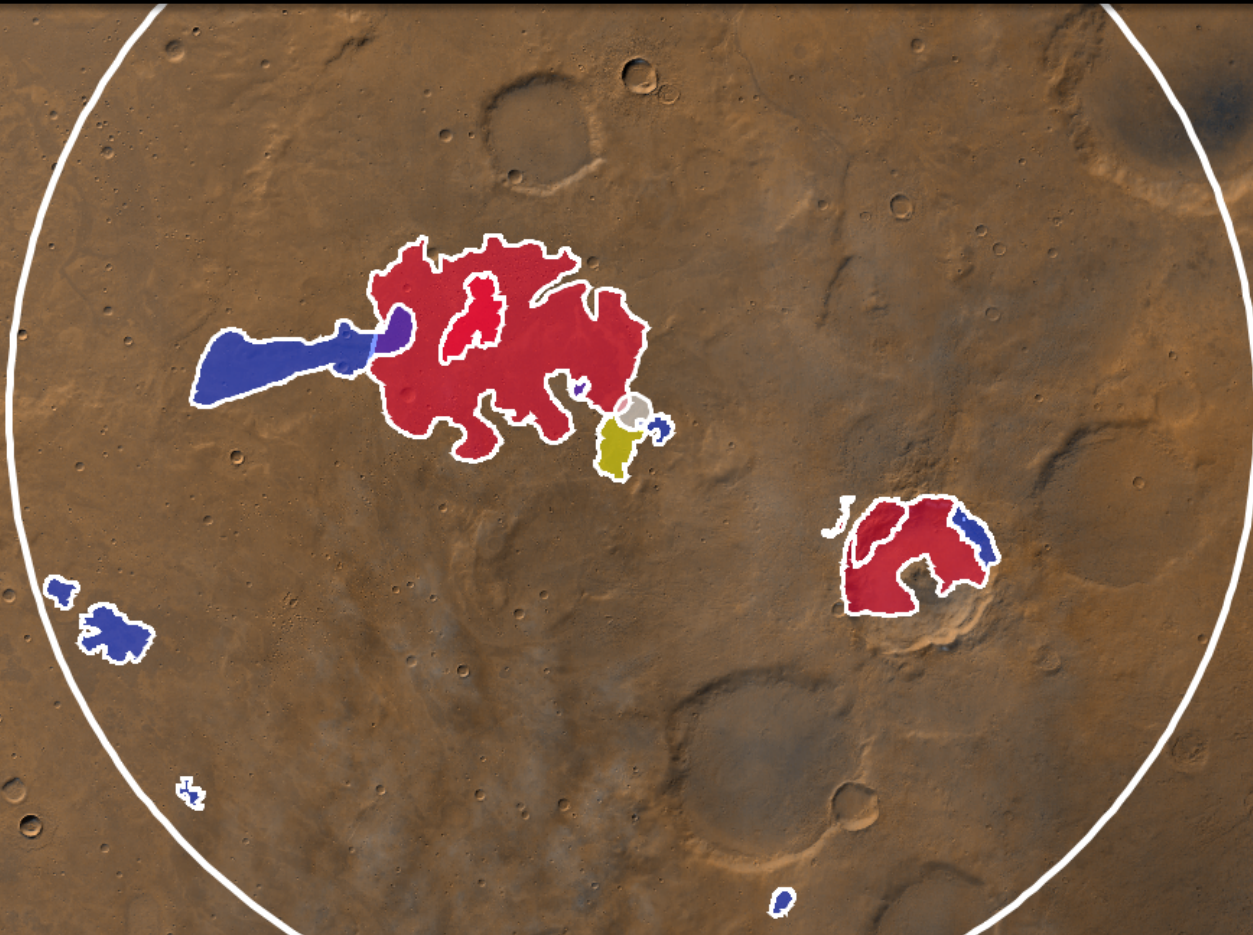


# Western Noachis Terra Chloride Deposits

Abstract #1021



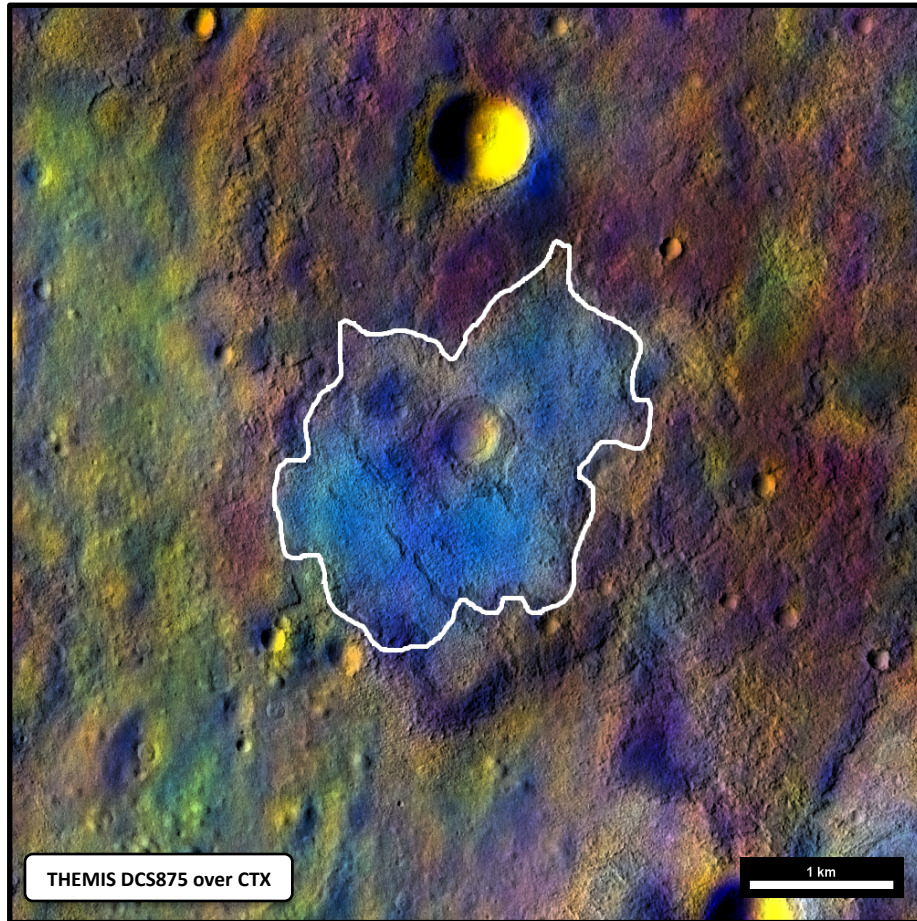
J. R. Hill<sup>1</sup> and P. R. Christensen<sup>1</sup>

<sup>1</sup>Arizona State University, Tempe, AZ 85287



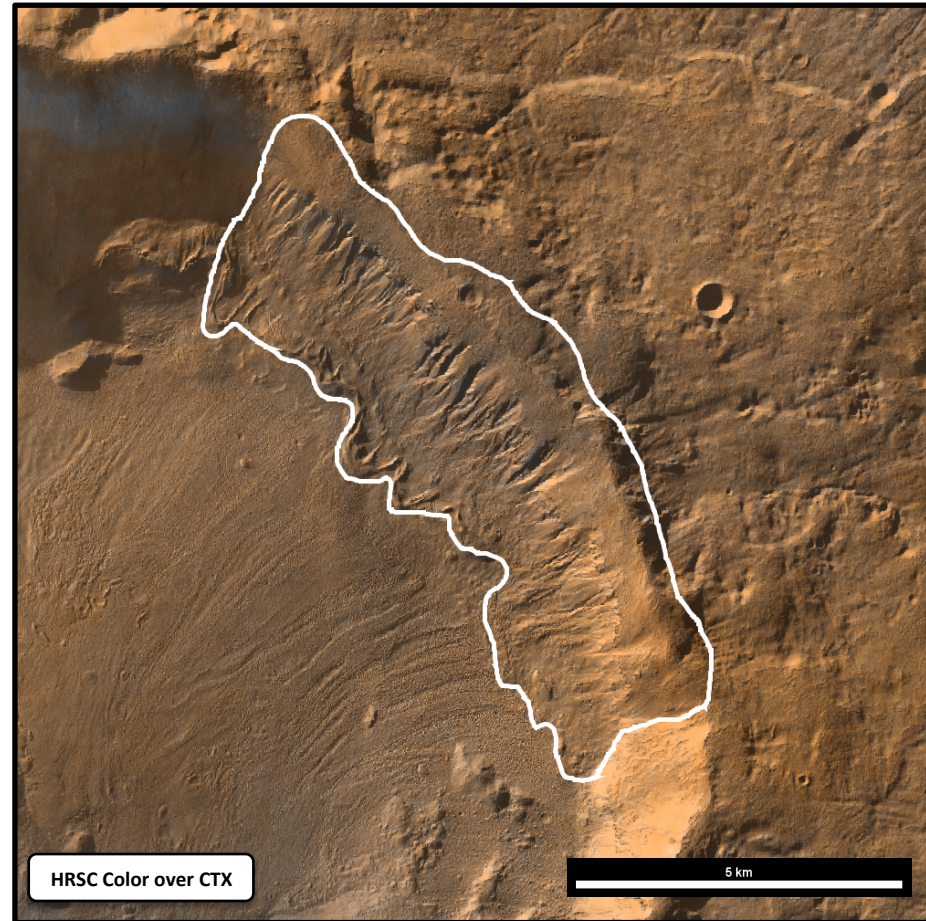
# Chlorides & Glacier-Like Forms (GLFs)

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Chloride Deposits

Based on *Osterloo et al., 2010*



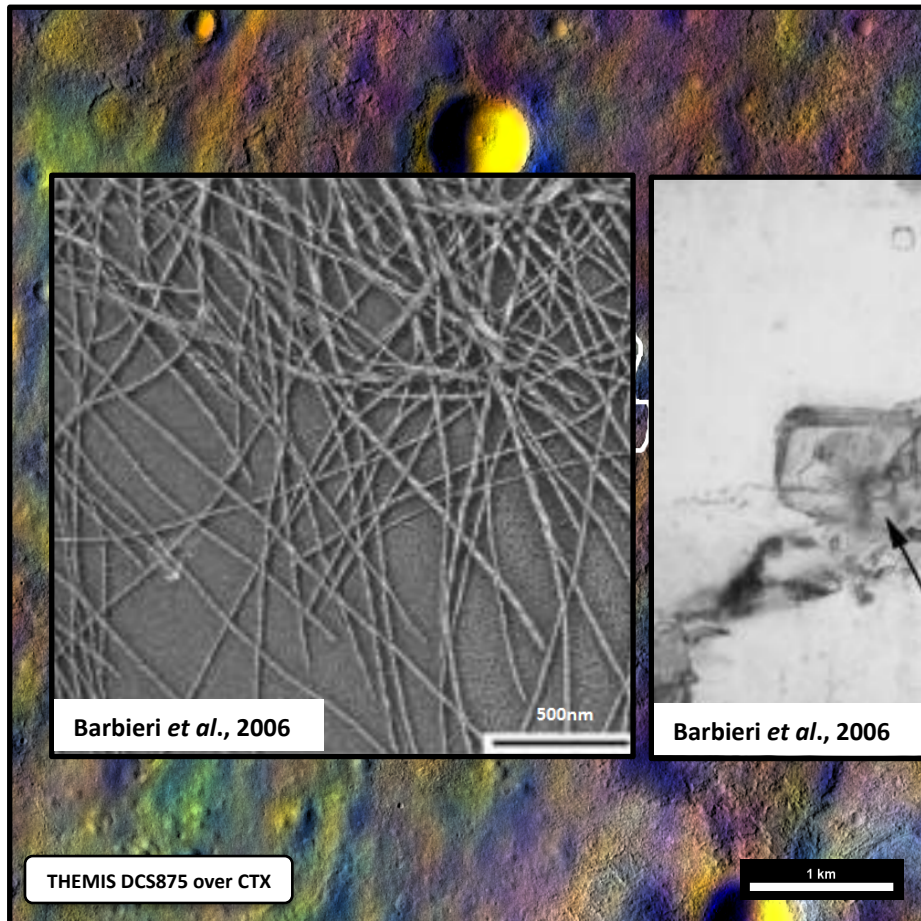
Glacier-Like Forms

Based on *Souness et al., 2012*



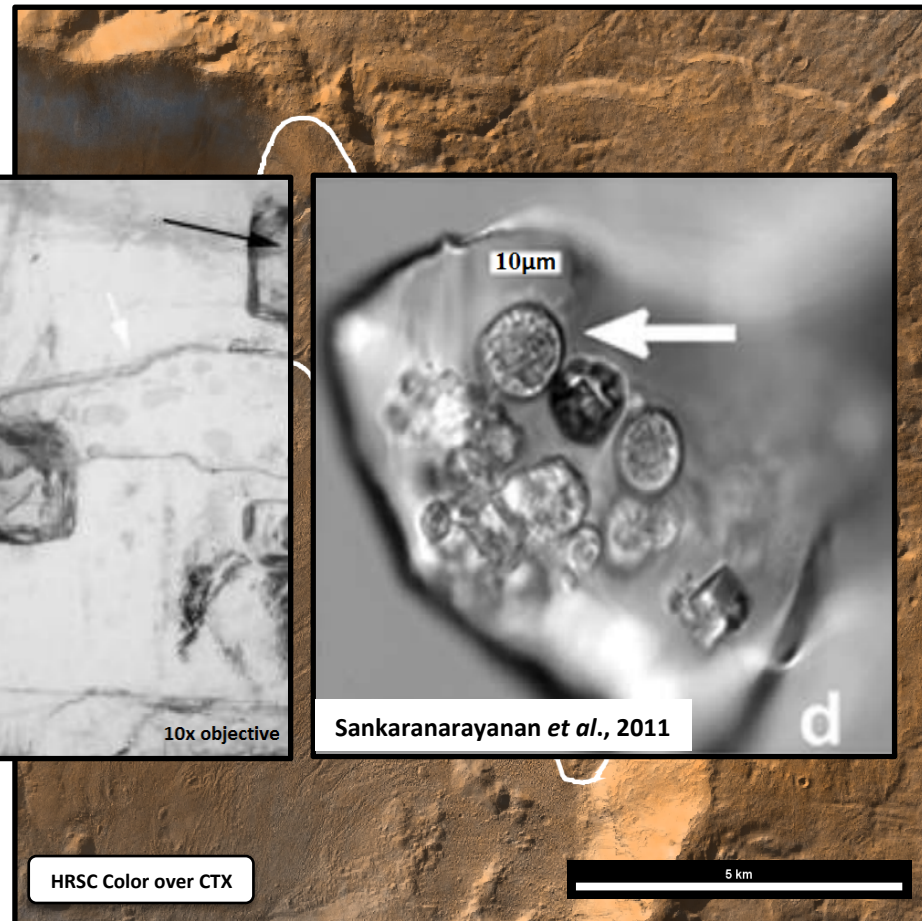
# Chlorides & Glacier-Like Forms (GLFs)

1<sup>st</sup> EZ Workshop for Human Missions to Mars



Chloride Deposits

Based on Osterloo et al., 2010



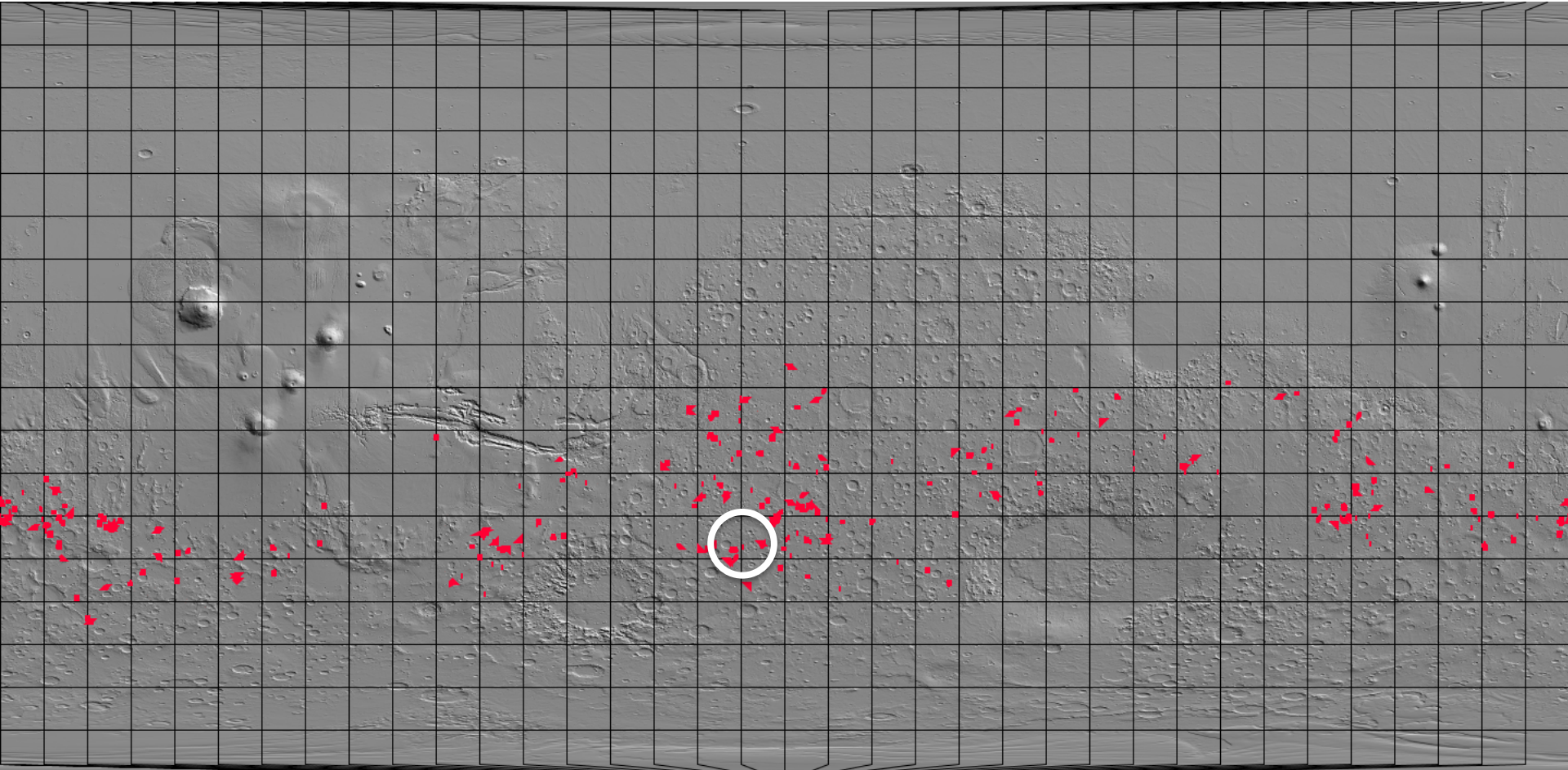
Glacier-Like Forms

Based on Souness et al., 2012

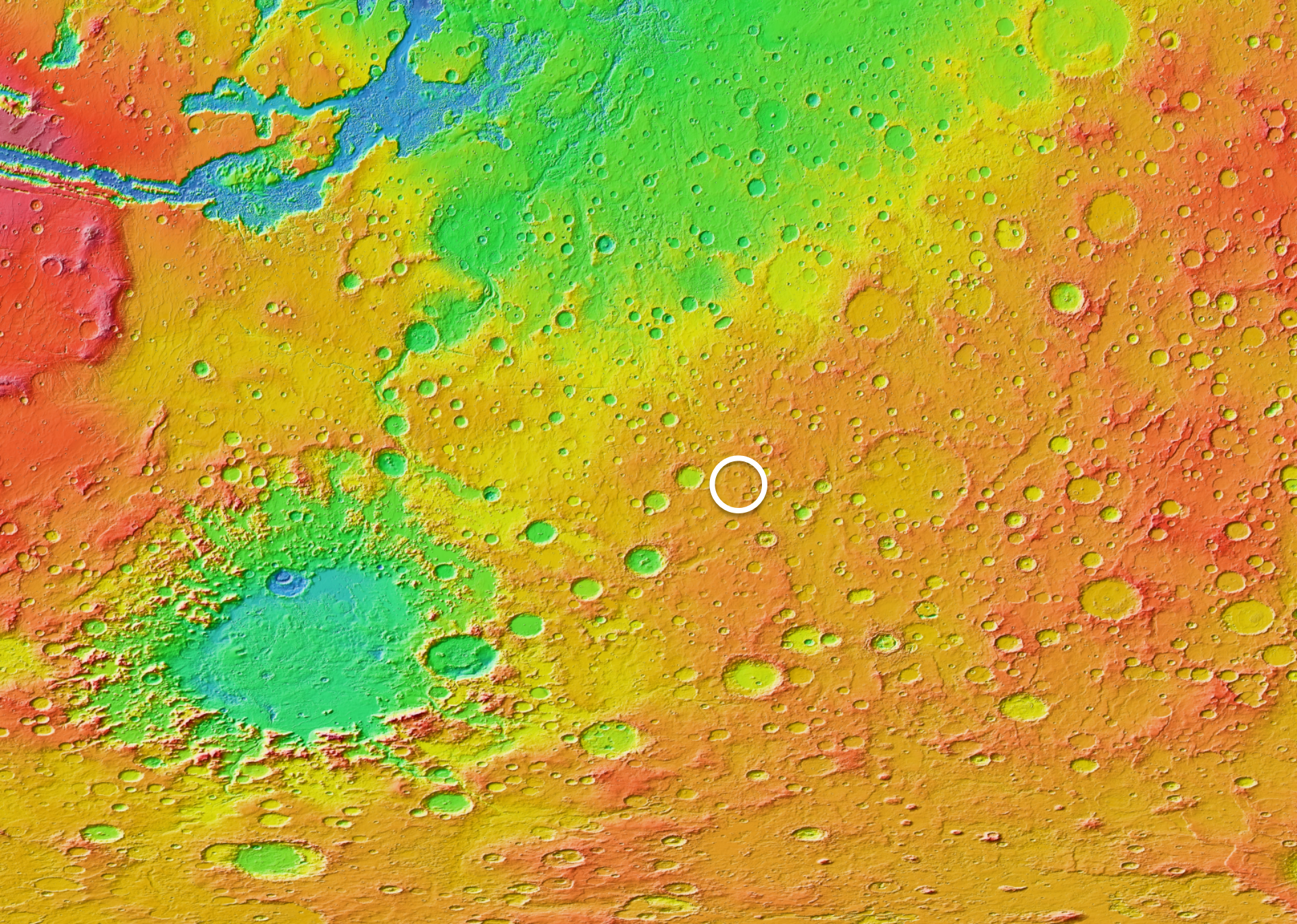


# Global Chloride Distribution

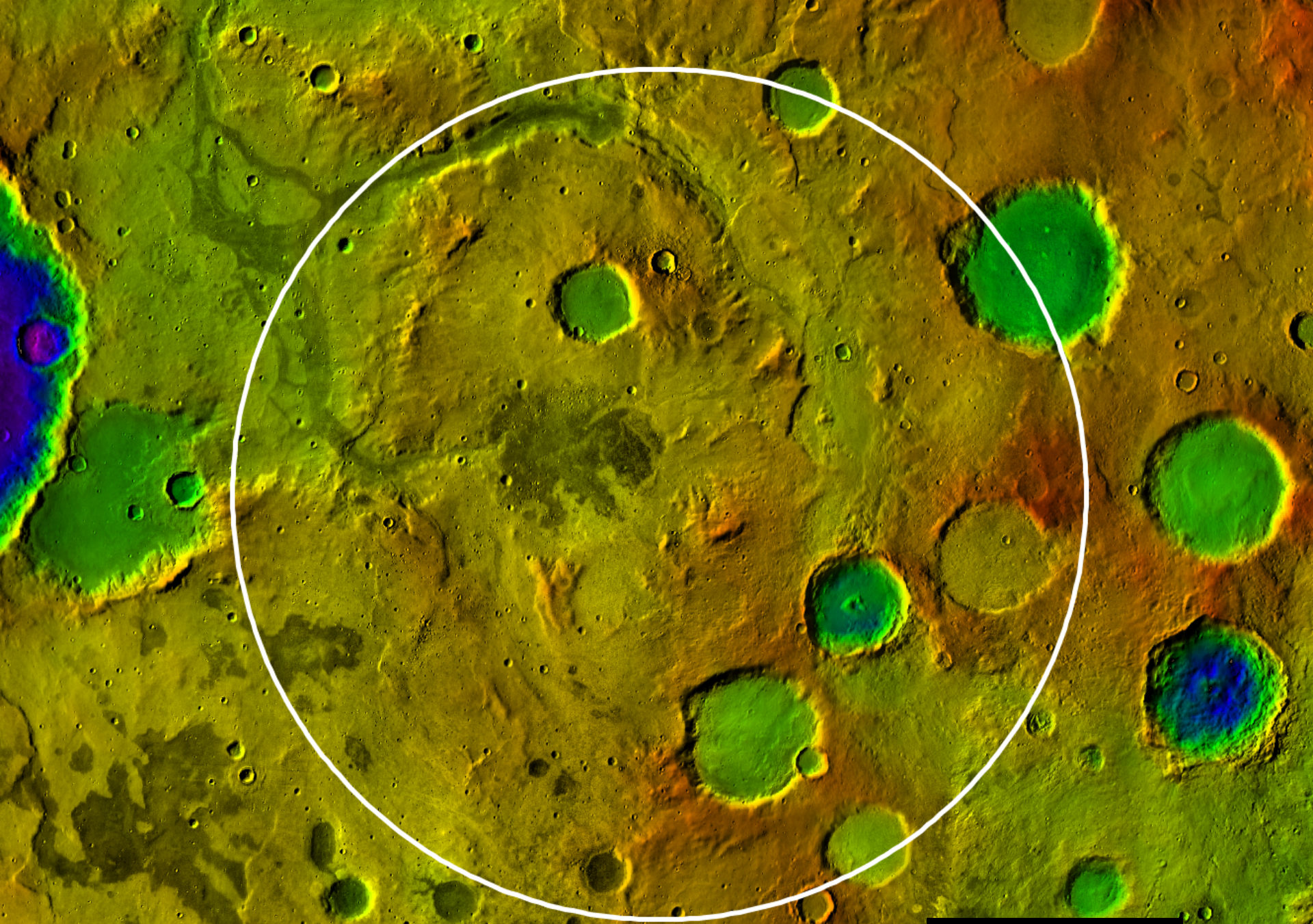
1<sup>st</sup> EZ Workshop for Human Missions to Mars











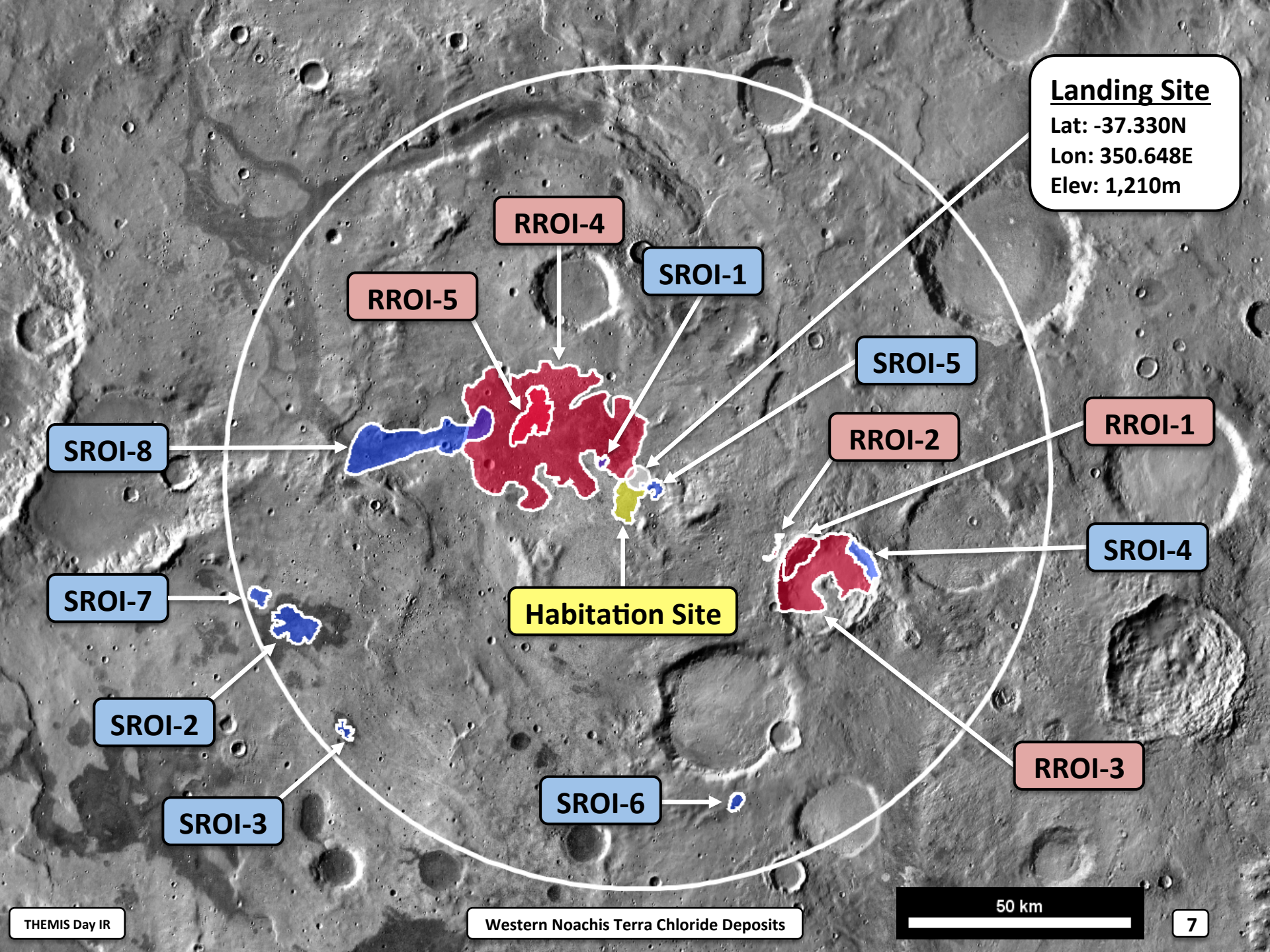
MOLA over THEMIS Day IR

Western Noachis Terra Chloride Deposits

50 km

6





**Landing Site**

Lat: -37.330N  
Lon: 350.648E  
Elev: 1,210m

RROI-4

SROI-1

RROI-5

SROI-5

SROI-8

RROI-2

RROI-1

SROI-7

Habitation Site

SROI-4

SROI-2

SROI-3

SROI-6

RROI-3



# Science ROIs

## Landing Site

Lat: -37.330N  
Lon: 350.648E  
Elev: 1,210m

SROI-8

SROI-1

SROI-5

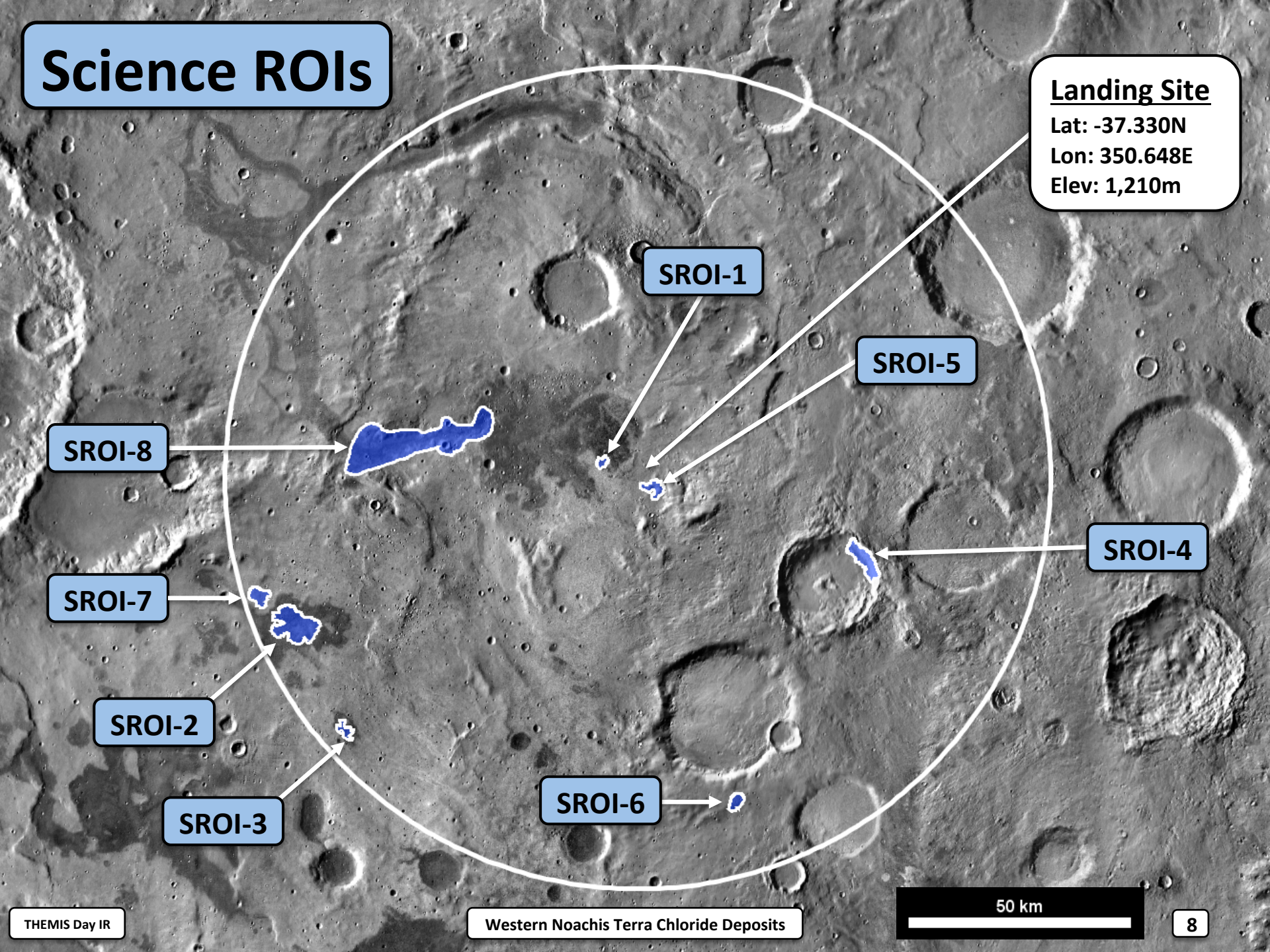
SROI-7

SROI-4

SROI-2

SROI-6

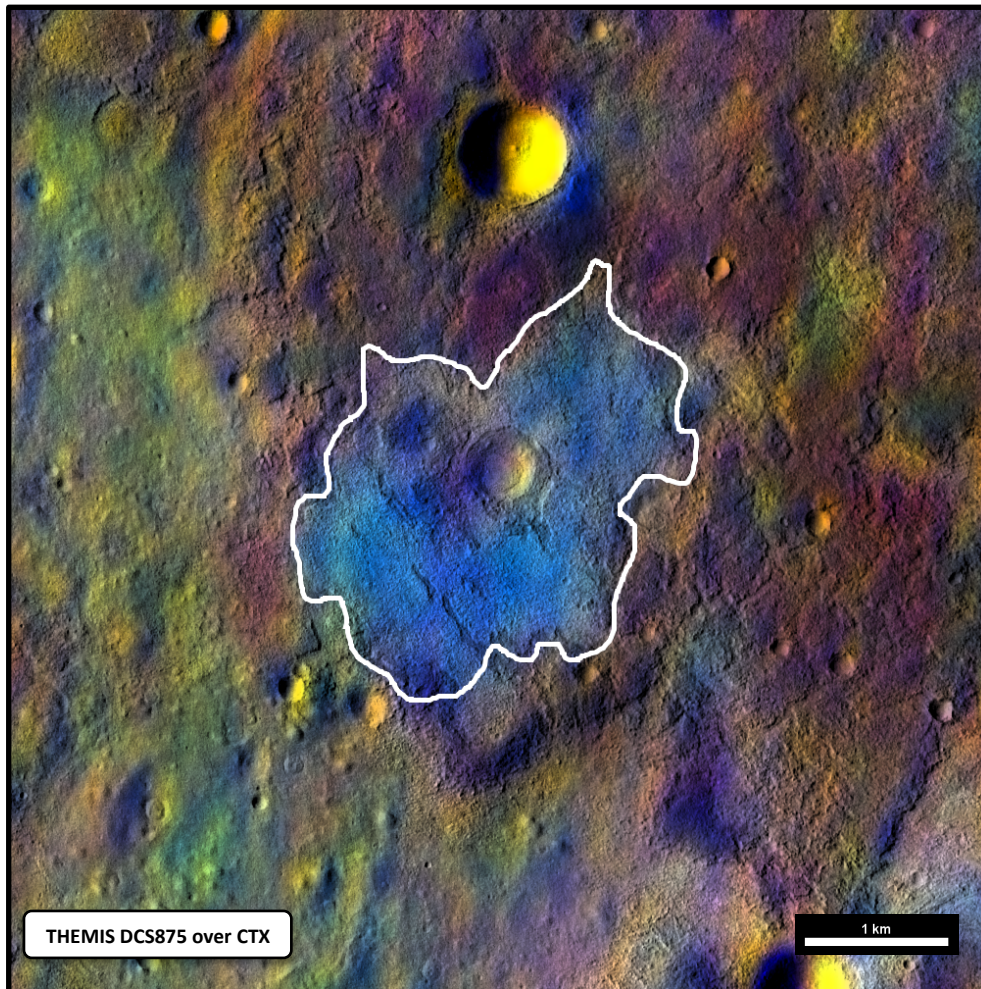
SROI-3





# Science ROI 1: Chloride Deposit 1

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.267N, 350.467E

**Elevation:** 1,158m

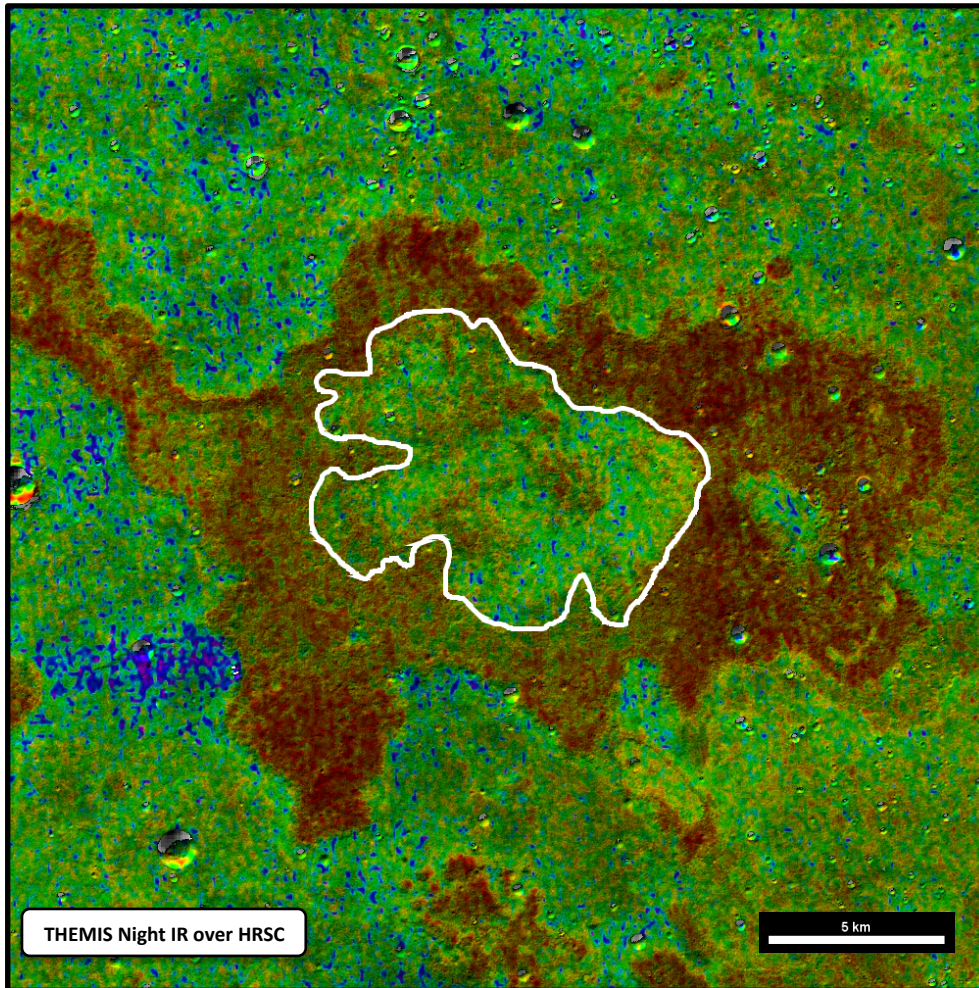
**Coverage:** THEMIS, CTX

- Chloride mineral(s) deposit with high astrobiological preservation potential
- Basaltic unit with ~15% olivine abundance (TES spectra), likely phyllosilicates, possibly hydrated minerals



# Science ROI 2: Chloride Deposit 2

1<sup>st</sup> EZ Workshop for Human Missions to Mars



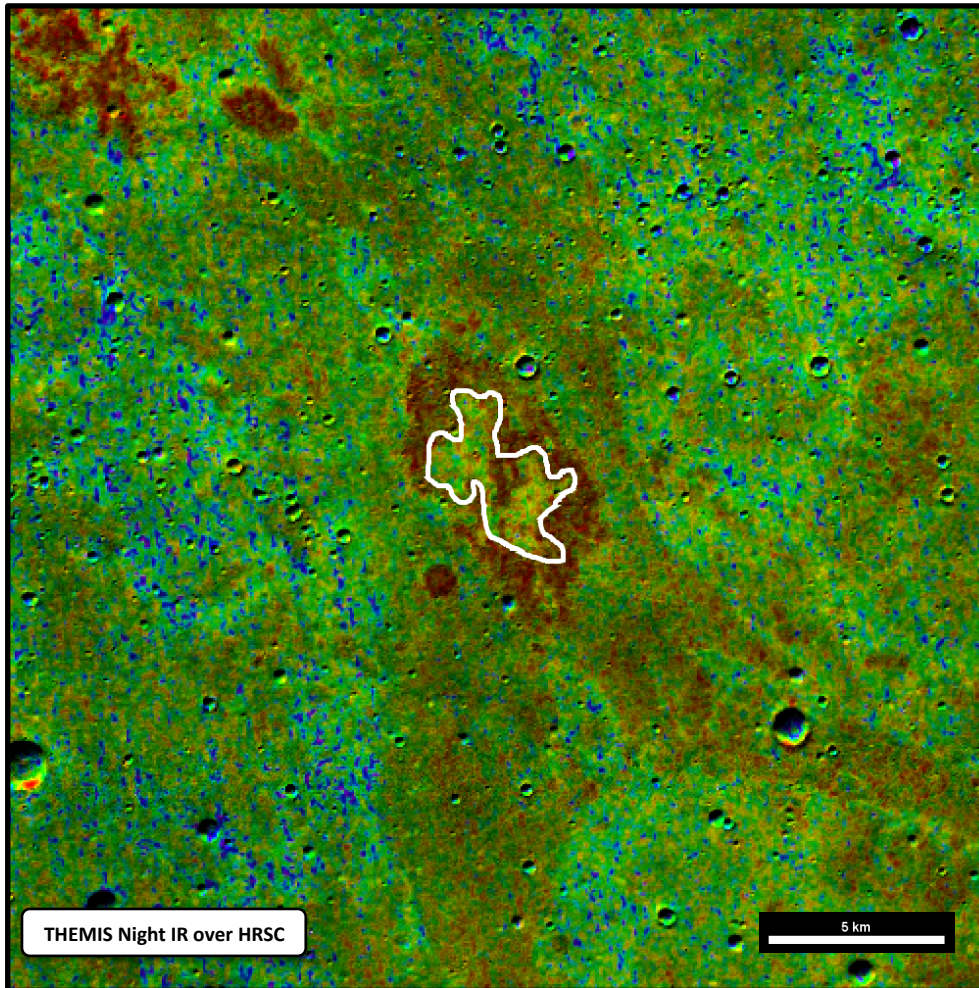
**Location:** -37.922N, 348.905E  
**Elevation:** 1,160m  
**Coverage:** THEMIS, CTX, HiRISE, CRISM

- Chloride mineral(s) deposit with high astrobiological preservation potential
- Basaltic unit with ~15% olivine abundance (TES spectra), phyllosilicates (CRISM) and possibly hydrated minerals (CRISM)



# Science ROI 3: Chloride Deposit 3

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -38.347N, 349.141E

**Elevation:** 1,192m

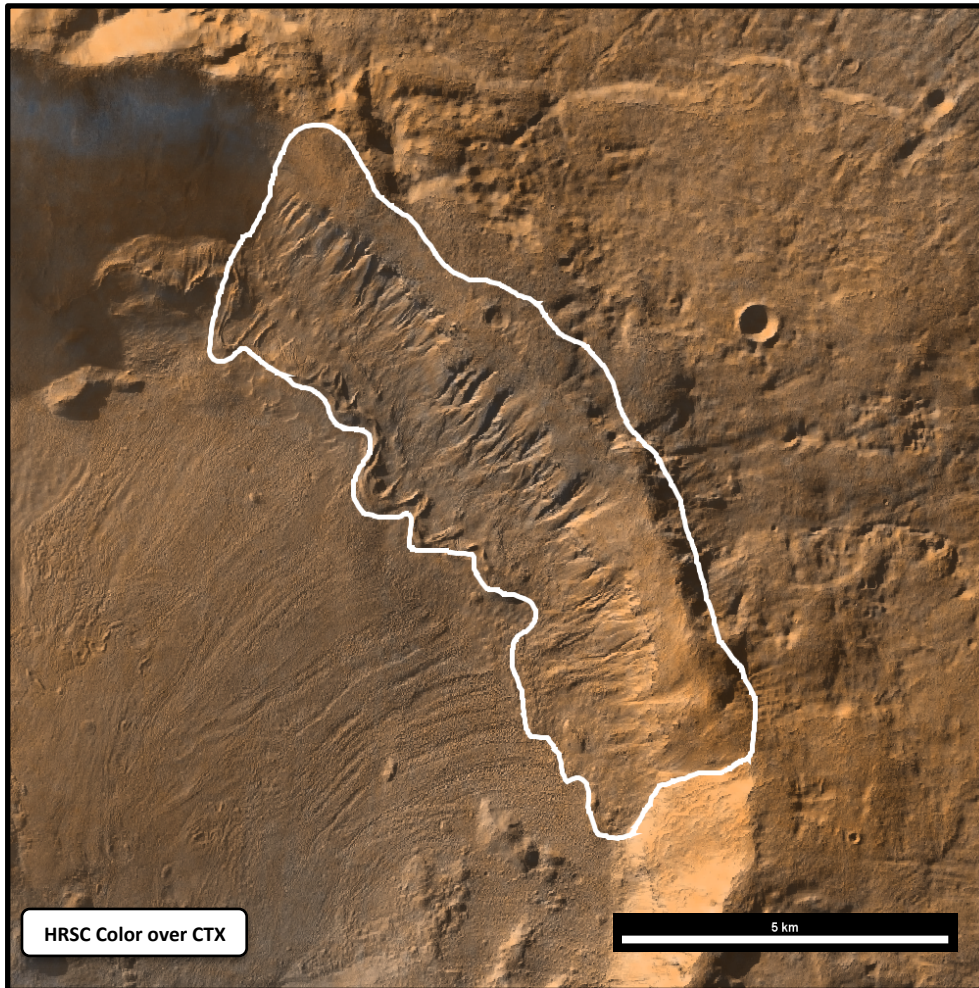
**Coverage:** THEMIS, CTX

- Chloride mineral(s) deposit with high astrobiological preservation potential
- Basaltic unit with ~15% olivine abundance (TES spectra), likely phyllosilicates, possibly hydrated minerals



# Science ROI 4: Gullies and Potential RSL

1<sup>st</sup> EZ Workshop for Human Missions to Mars



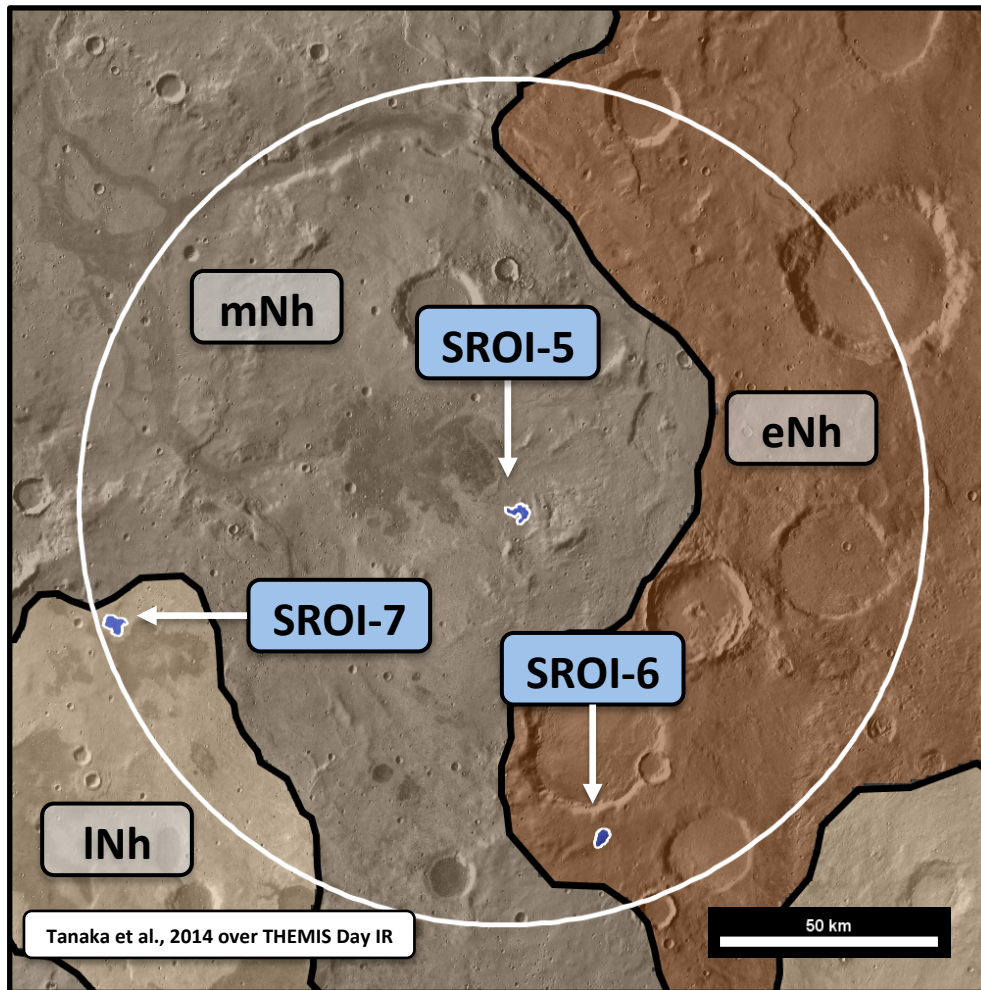
**Location:** -37.823N, 351.823E  
**Elevation:** 668m – 1,711m  
**Coverage:** THEMIS, CTX, HiRISE, CRISM

- Gullies and glacier-like forms on northeastern crater rim
- Morphology similar to RSL-bearing crater rims
  - If RSLs, access to potentially habitable environment
  - If not, access to an RSL analog environment without planetary protection concerns



# Science ROIs 5,6, & 7

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**eNh = Early Noachian  
Highlands**

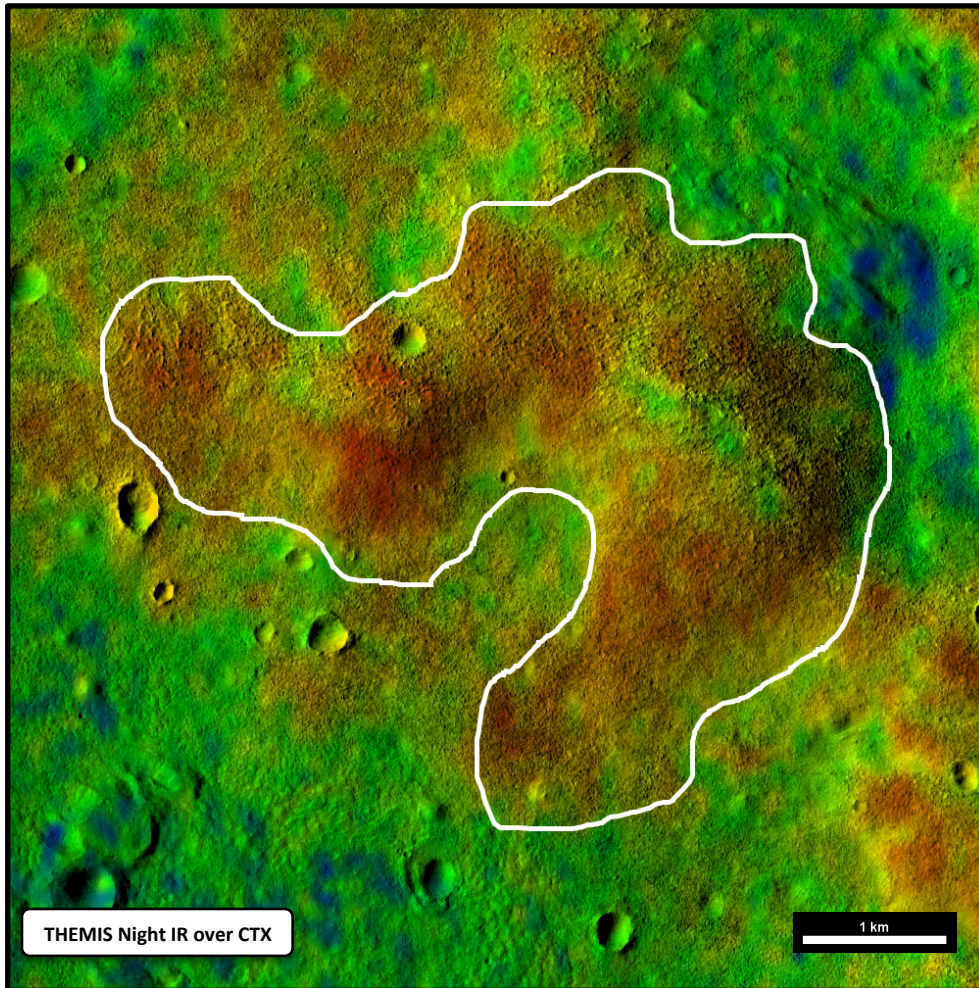
**mNh = Middle Noachian  
Highlands**

**INh = Late Noachian  
Highlands**



# Science ROI 5: Middle Noachian Material

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.368N, 350.732E

**Elevation:** 1,231m – 1,389m

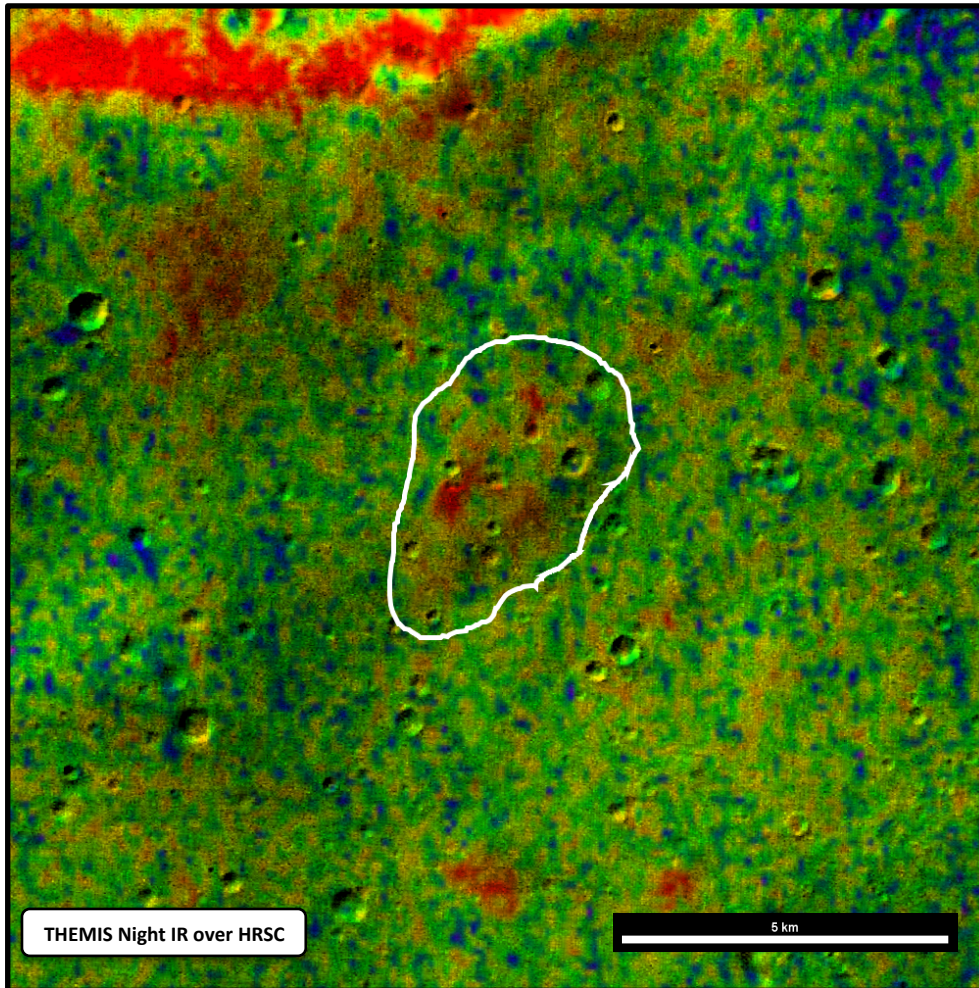
**Coverage:** THEMIS, CTX

- High thermal inertia region with potential outcrops of middle Noachian material
- Lies partially within the proposed landing site



# Science ROI 6: Early Noachian Material

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -38.667N, 351.165E

**Elevation:** 1,546m – 1,655m

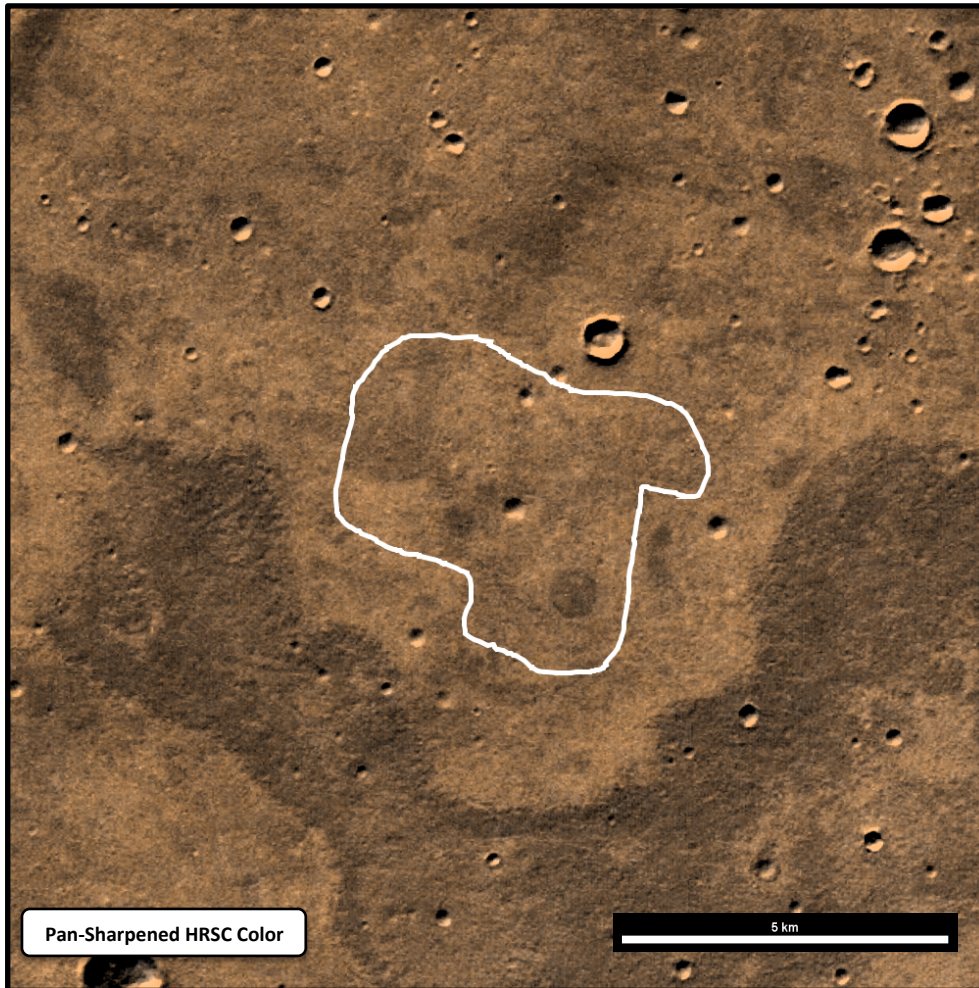
**Coverage:** THEMIS, HRSC

- Slightly higher thermal inertia region with potential outcrops of early Noachian material
- Remnant (negative) magnetic field based on MAG/ER measurements



# Science ROI 7: Late Noachian Material

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.817N, 348.699E

**Elevation:** 1,192m – 1,217m

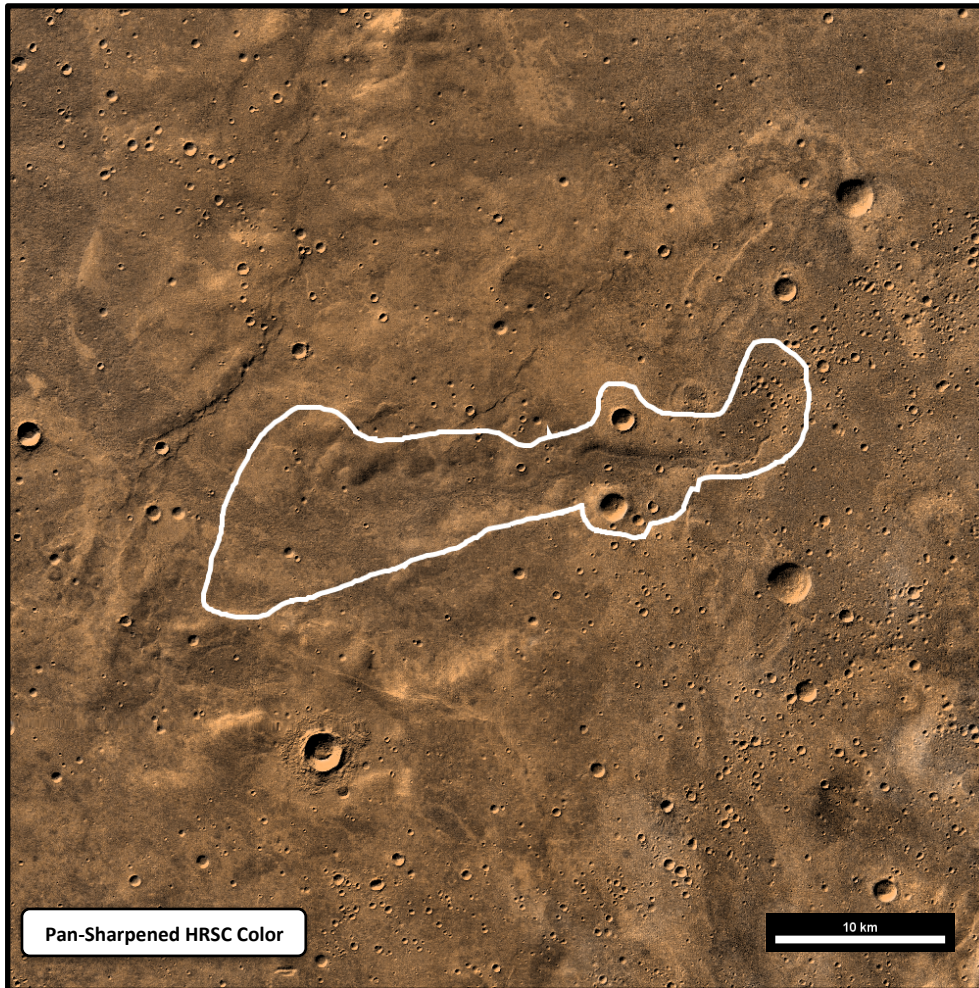
**Coverage:** THEMIS, HRSC

- Slightly higher thermal inertia region with potential outcrops of late Noachian material
- Remnant (positive) magnetic field based on MAG/ER measurements



# Science ROI 8: Channel with Stratigraphy

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.184N, 349.537E

**Elevation:** 1,105m – 1,204m

**Coverage:** THEMIS, HRSC

- Outflow channel from the basin containing the primary chloride deposit
- Stratigraphy likely exposed in the steep-walled sections of the channel



# Resource ROIs

## Landing Site

Lat: -37.330N  
Lon: 350.648E  
Elev: 1,210m

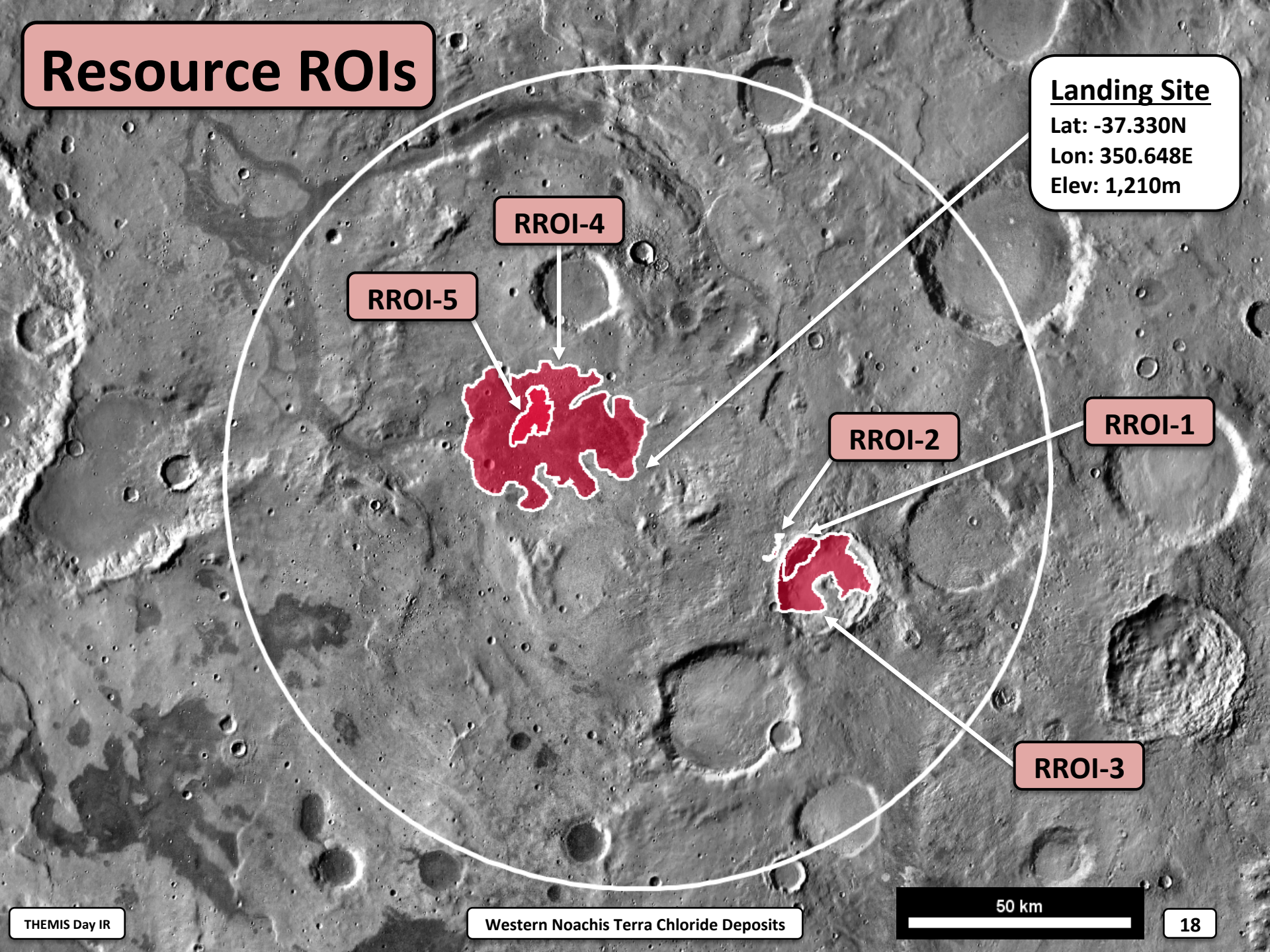
RROI-4

RROI-5

RROI-2

RROI-1

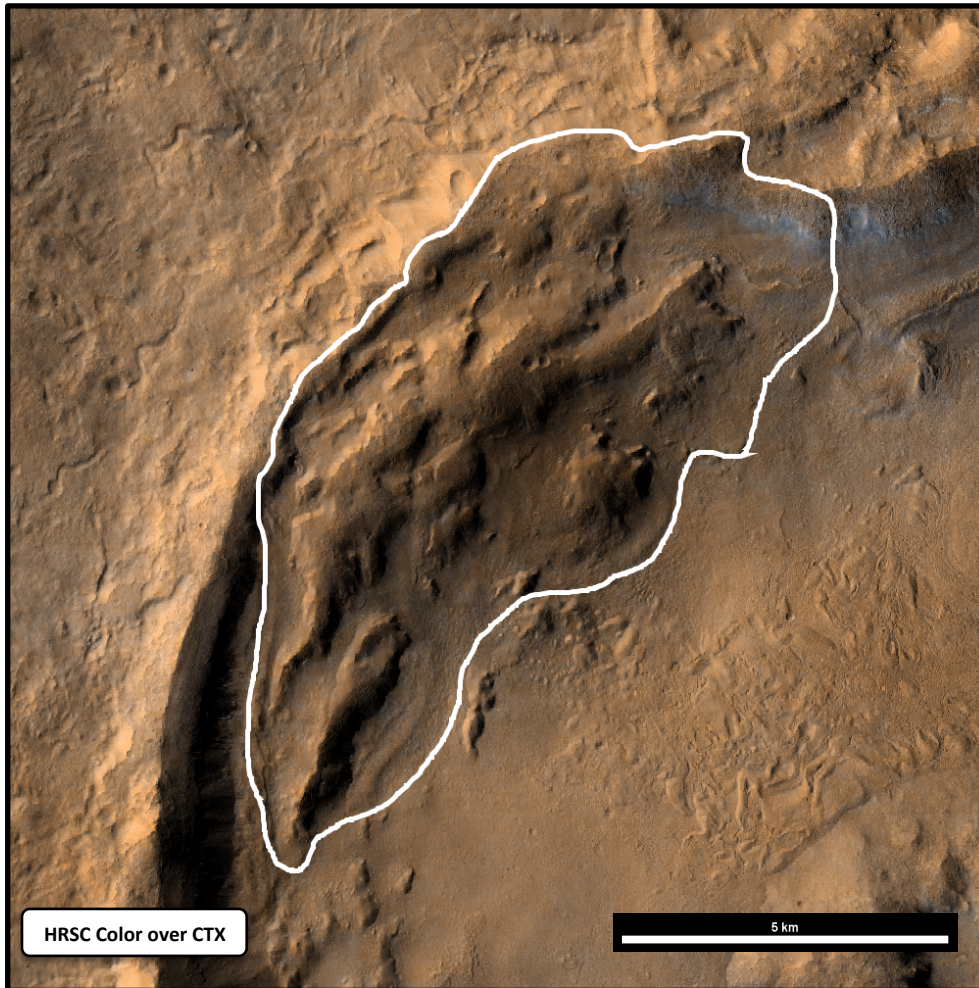
RROI-3





# Resource ROI 1: Pasted Terrain

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.636N, 351.475E

**Elevation:** 825m – 1,476m

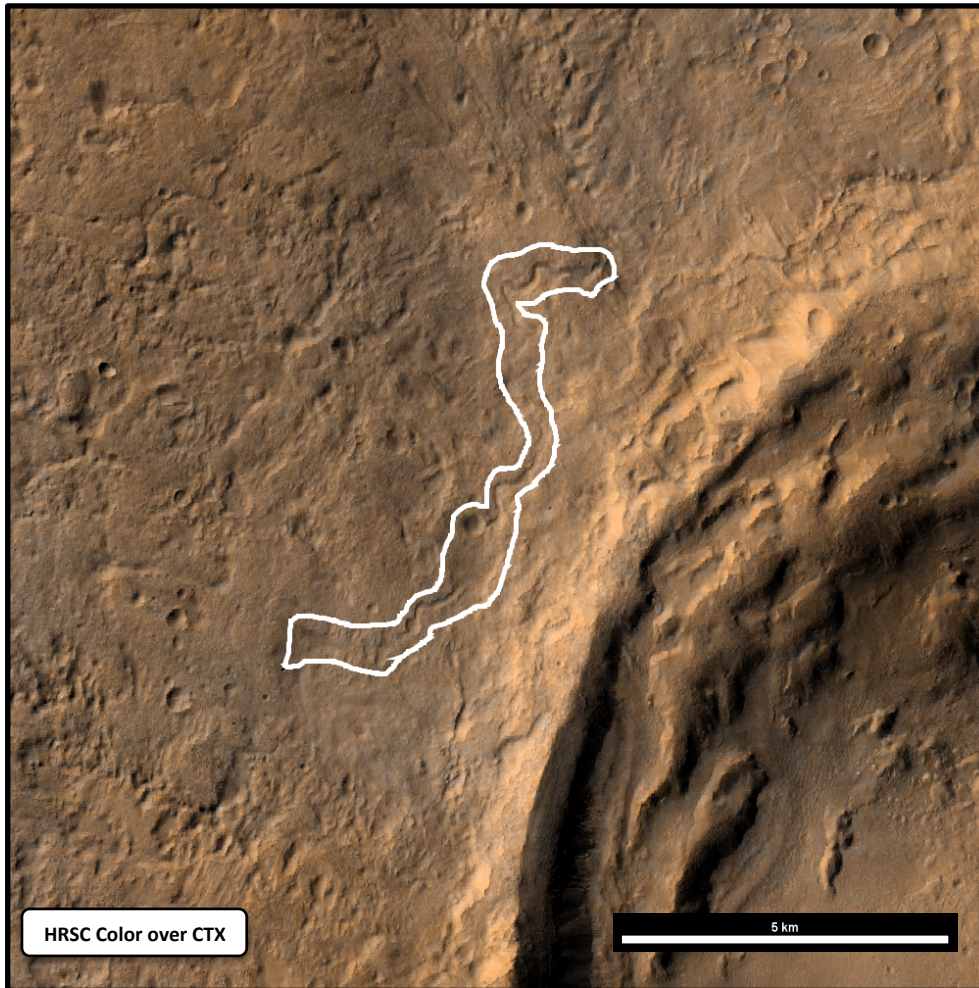
**Coverage:** THEMIS, HRSC, CTX

- Residual ice under thin armor of dust/regolith on south-facing crater rim
- Patches closer to the rim are more accessible from the landing site



# Resource ROI 2: Lineated Valley Fill

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.616N, 351.367E

**Elevation:** 1,247m – 1,355m

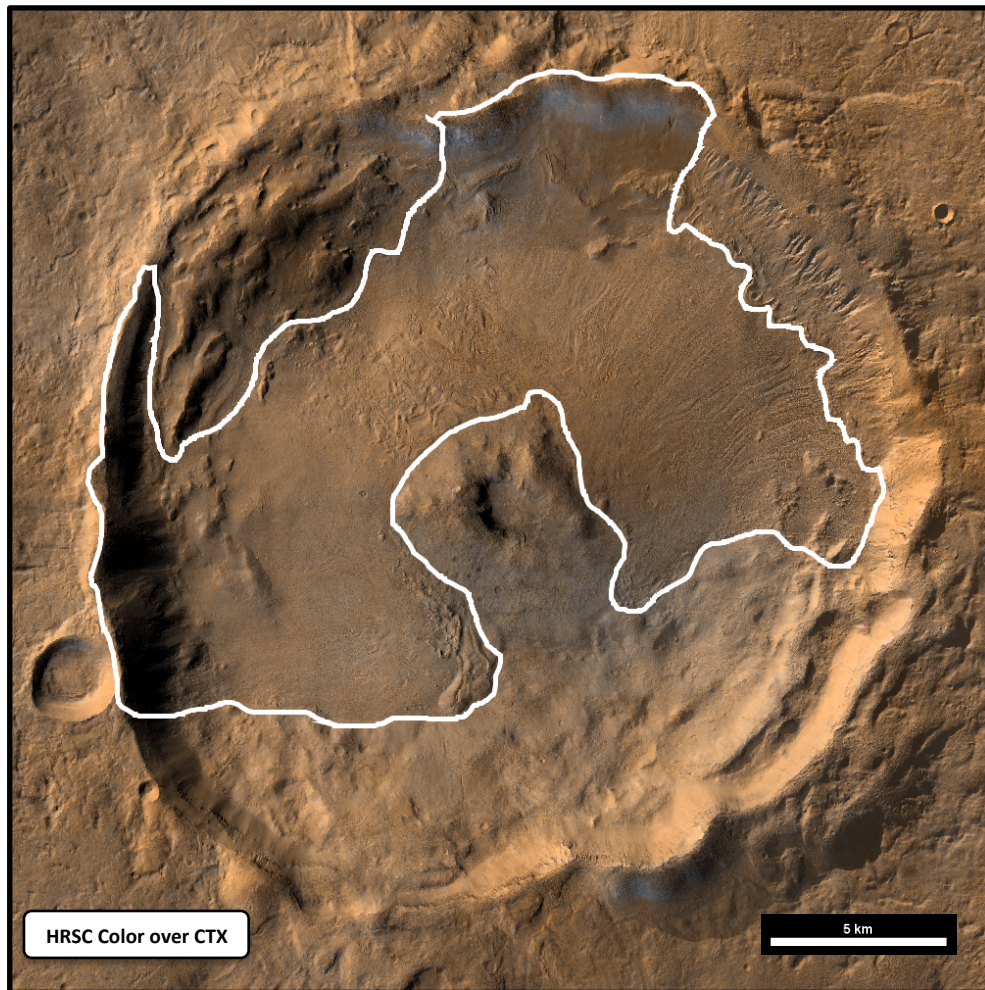
**Coverage:** THEMIS, HRSC, CTX

- Potential lineated valley fill in crater ejecta
- Could indicate additional subsurface water ice deposits
- More accessible than RROI-1, but more difficult to confirm due to lack of high resolution data



# Resource ROI 3: Lineated Crater Fill

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.753N, 351.625E

**Elevation:** 273m – 1,731m

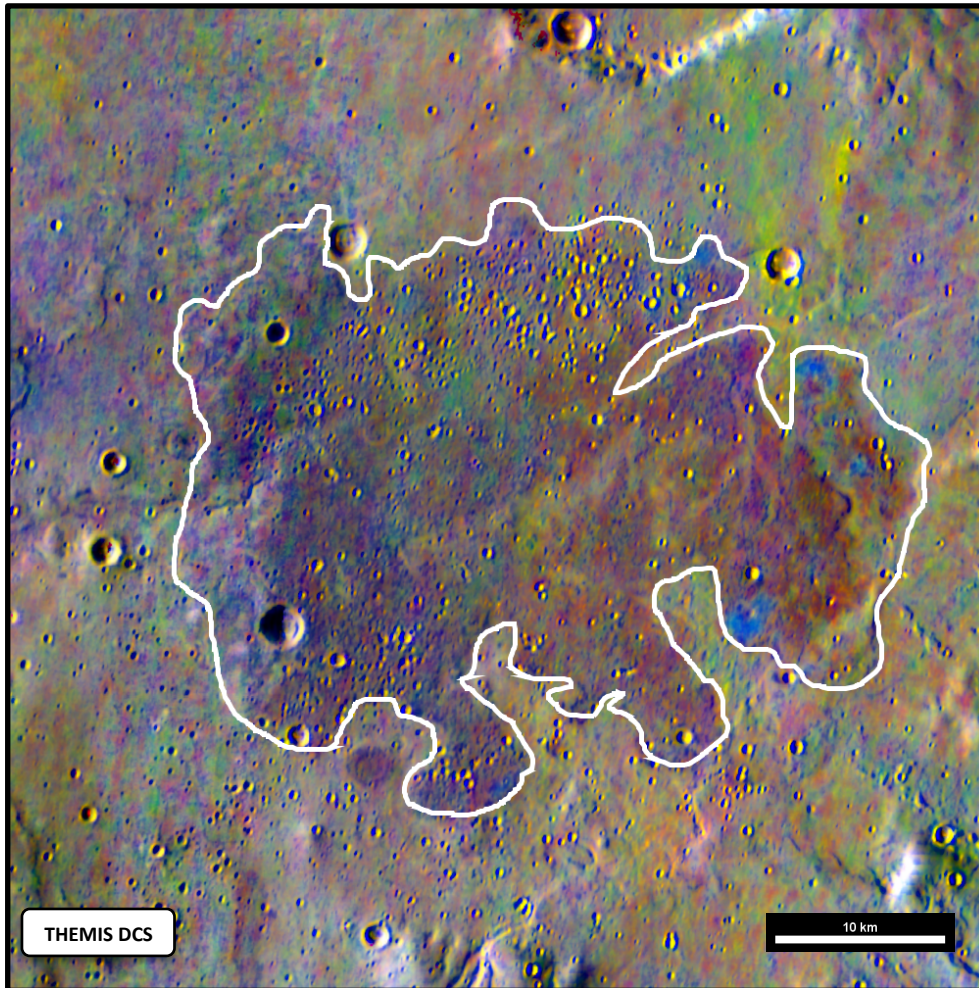
**Coverage:** THEMIS, HRSC, CTX

- Subsurface ice under thin armor of dust/regolith on crater floor
- Appears similar to concentric crater fill, but is not continuous around the crater
- SHARAD profiles have been acquired over this deposit, but need analysis



# Resource ROI 4: Olivine-Bearing Basalt

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.141N, 350.188E

**Elevation:** 1,133m – 1,213m

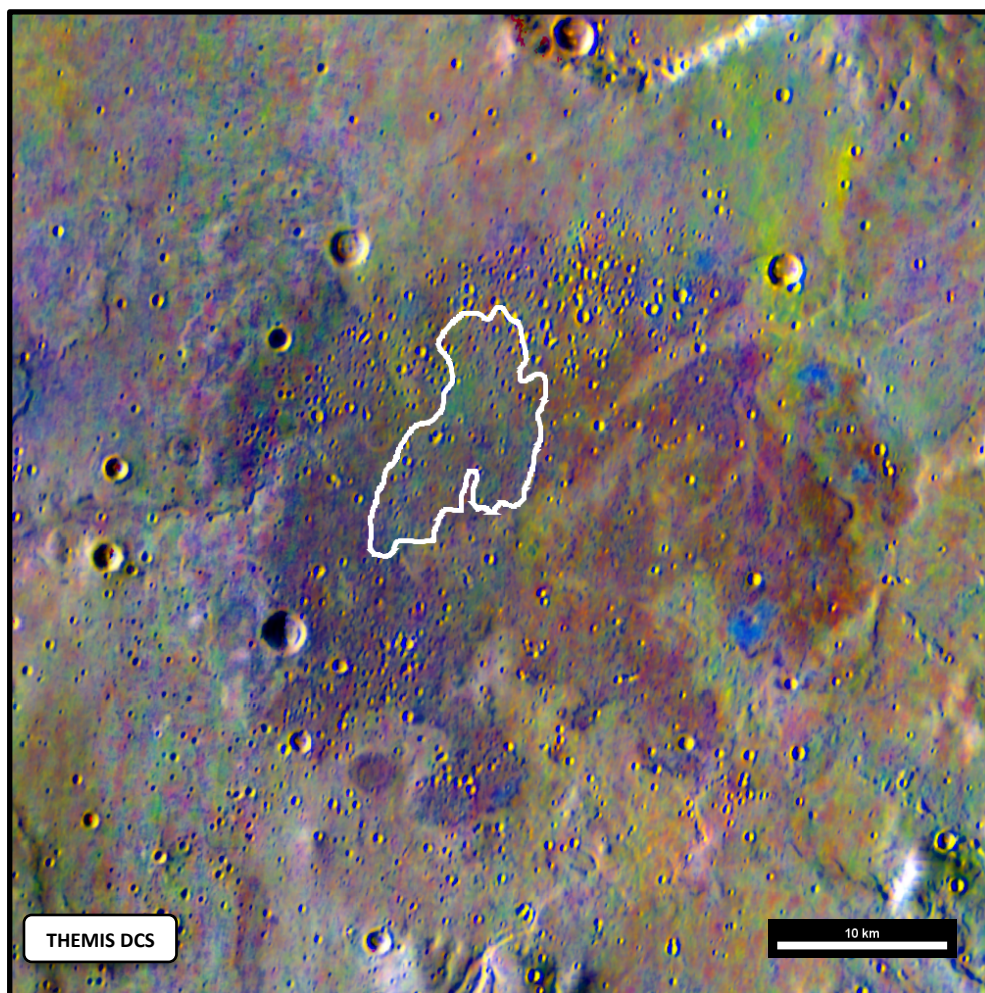
**Coverage:** THEMIS, HRSC, CTX

- Basaltic volcanic unit that appears to underlie the primary chloride deposit
- TES spectra show an ~15% olivine abundance, but Fo number not yet determined
- CRISM data expected to confirm presence of hydrated minerals



# Resource ROI 5: Silica-Enriched Material

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.072N, 350.120E

**Elevation:** 1,130m – 1,157m

**Coverage:** THEMIS, HRSC, CTX

- Additional region of silica enrichment on the surface
- TES spectra indicate an ~25% abundance of a high-silica material



# Science ROIs Rubric

1<sup>st</sup> EZ Workshop for Human Missions to Mars

Site Factors					SROI-1	SROI-2	SROI-3	SROI-4	SROI-5	SROI-6	SROI-7	SROI-8	RROI-1	RROI-2	RROI-3	RROI-4	RROI-5	LS	HS	EZ SUM
Science Site Criteria	Astrobio	Threshold	AND/OR	Potential for past habitability	●	●	●													3,0
				Potential for present habitability/refugia				○												0,1
		Qualifying		Potential for organic matter, w/ surface exposure	●	●	●													3,0
	Atmospheric Science	Threshold		Noachian/Hesperian rocks w/ trapped atmospheric gases	○	○	○		○	○	○	○				○	○	○	○	0,11
		Qualifying		Meteorological diversity in space and time	●	●	●	●				●	●	●	●		○			8,1
				High likelihood of surface-atmosphere exchange				○					○	○	○					0,4
				Amazonian subsurface or high-latitude ice or sediment				●					●	●	●					4,0
				High likelihood of active trace gas sources	?	?	?	?	?	?	?	?	?	?	?	?	?	?	?	?
	Geoscience	Threshold		Range of martian geologic time; datable surfaces	○	○	○		●	●	●				○	○	○	●	●	5,6
				Evidence of aqueous processes	●	●	●	●				●		●						6,0
				Potential for interpreting relative ages	●	●	●	●				●				●	●			8,0
		Qualifying		Igneous Rocks tied to 1+ provinces or different times	●	●	●		●	●	●	●				●	●	●	○	10,0
				Near-surface ice, glacial or permafrost				○					●	○	●					2,2
				Noachian or pre-Noachian bedrock units	○	○	○		●	●	●					○	○	●	●	5,5
				Outcrops with remnant magnetization						●	●									2,0
				Primary, secondary, and basin-forming impact deposits	○	○	○							●		○	○			1,5
				Structural features with regional or global context																0,0
				Diversity of aeolian sediments and/or landforms	○	○	○									●				1,3



# Resource ROIs Rubric

1<sup>st</sup> EZ Workshop for Human Missions to Mars

Site Factors			SROI-1	SROI-2	SROI-3	SROI-4	SROI-5	SROI-6	SROI-7	SROI-8	RROI-1	RROI-2	RROI-3	RROI-4	RROI-5	LS	HS	EZ SUM
ISRU and Civil Engineering Criteria	Engineering	Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)		●	●	●	●	●	●	●	●	●	●	●	●	●	●	15,0
	Water Resource	Threshold	AND/OR	Potential for ice or ice/regolith mix		●					●	○	●					3,1
				Potential for hydrated minerals		○	○	○						○	○			0,5
			Quantity for substantial production			●					●	○	●					3,1
			Potential to be minable by highly automated systems			○					●	○	●					2,2
			Located less than 3 km from processing equipment site		○													0,1
			Located no more than 3 meters below the surface			○					○	○	○					0,4
			Accessible by automated systems		●	●	●	○			○	○	○	●	●			5,4
		Qualifying	Potential for multiple sources of ice, ice/regolith mix <b>and</b> hydrated minerals		○	●	○	●			●	○	●	○	○			4,5
			Distance to resource location can be >5 km		○	●	●	●			●	●	●	●	●			9,1
			Route to resource location must be (plausibly) traversable		●	●	●	○			●	●	○	●	●			7,2
	Civil Engineering	Threshold	~50 sq km region of flat and stable terrain with sparse rock distribution														●	1,0
			1–10 km length scale: <10°													●	●	2,0
			Located within 5 km of landing site location														●	1,0
		Qualifying	Located in the northern hemisphere															0,0
			Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith		●	●	●	○	○	○	○			●	●	○	●	6,5
	Food Production	Qualifying	Utilitarian terrain features														○	0,1
			Low latitude															0,0
			No local terrain feature(s) that could shadow light collection facilities		●	●	●		●	●	●	○	○	●	●	●	●	10,2
			Access to water		○	●	○	●			●	○	●	○	○			4,5
	Metal/Silicon Resource	Threshold	Access to dark, minimally altered basaltic sands		●	●	●		○	○	○	○		●	●	○		5,5
			Potential for metal/silicon		●	●	●							●	●			5,0
			Potential to be minable by highly automated systems		●	●	●							●	●			5,0
			Located less than 3 km from processing equipment site		●													1,0
			Located no more than 3 meters below the surface		●	●	●							●	●			5,0
			Accessible by automated systems		●	●	●							○	●			4,1
		Qualifying	Potential for multiple sources of metals/silicon		●	●	●							●	●			5,0
			Distance to resource location can be >5 km		●	●	●							●	●			5,0
			Route to resource location must be (plausibly) traversable		●	●	●							●	○			4,1



# Highest Priority EZ Data Needs

1<sup>st</sup> EZ Workshop for Human Missions to Mars



## Current Assets

- **Science:** HiRISE & CRISM FRT Pair of SROI-1
- **Resource:** CTX Stereo Pair of Traverse Route Between LS & RROI-1/RRO1-2

## Potential Future Assets

- **Science:** Thermal IR Spectra  $>25\mu\text{m}$
- **Resource:** Higher Resolution Subsurface Radar



# Summary

1<sup>st</sup> EZ Workshop for Human Missions to Mars

## ■ Chloride Deposits

- Remnants of potentially habitable environments
- High astrobiological preservation potential

## ■ Near-Surface Water Ice

- Significant water resources
- Amazonian climate science

## ■ Additional ROIs representing a variety of other science and resource targets





# Backup Slides



# References

1<sup>st</sup> EZ Workshop for Human Missions to Mars

- Barbieri, Roberto, et al. "Microbial signatures in sabkha evaporite deposits of Chott el Gharsa (Tunisia) and their astrobiological implications." *Planetary and Space Science* 54.8 (2006): 726-736.
- Osterloo, M. M., F. S. Anderson, V. E. Hamilton, and B. M. Hynek (2010), Geologic context of proposed chloride bearing materials on Mars, *J. Geophys. Res.*, 115, E10012, doi: 10.1029/2010JE003613.
- Sankaranarayanan K, Timofeeff MN, Spathis R, Lowenstein TK, Lum JK (2011) Ancient Microbes from Halite Fluid Inclusions: Optimized Surface Sterilization and DNA Extraction. *PLoS ONE* 6(6): e20683. doi:10.1371/journal.pone.0020683.
- Souness, Colin, et al. "An inventory and population-scale analysis of martian glacier-like forms." *Icarus* 217.1 (2012): 243-255.



# Prioritization List of EZ Data Needs

1<sup>st</sup> EZ Workshop for Human Missions to Mars



## Science

- 1) HiRISE & CRISM FRT Pair of SROI-1
  - Verify stratigraphic relationship with volcanic unit and confirm co-occurrence of phyllosilicates
- 2) HiRISE & CRISM FRT Pair of SROI-3
  - Verify stratigraphic relationship with volcanic unit and confirm co-occurrence of phyllosilicates

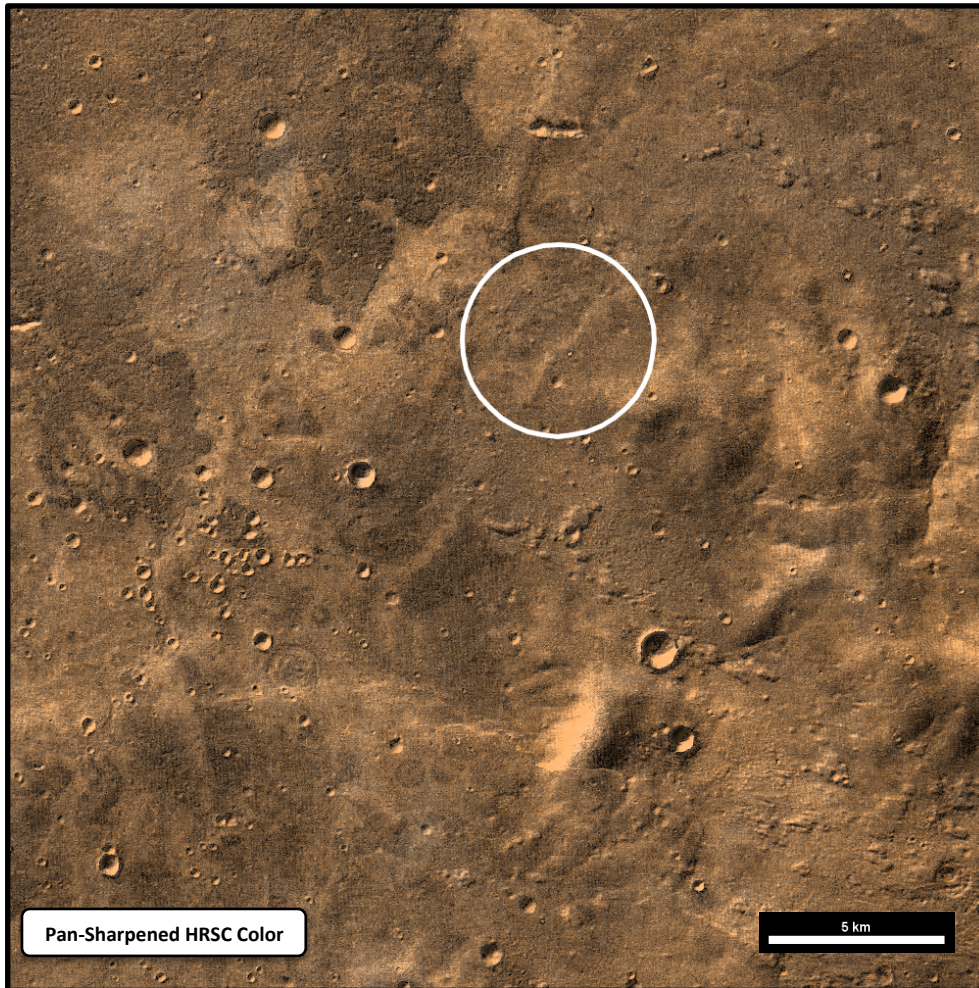
## Resource

- 1) CTX Stereo Pair of Traverse Route Between LS & RROI-1/2
  - Verify plausible traverse route
- 2) HiRISE & CRISM FRT Pair of RROI-2
  - Verify presence of lineated valley fill
- 3) HiRISE image(s) of Traverse Route Between LS & RROI-1/2
  - Plot plausible traverse route



# Landing Site

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.330N, 350.648E

**Elevation:** 1,210m

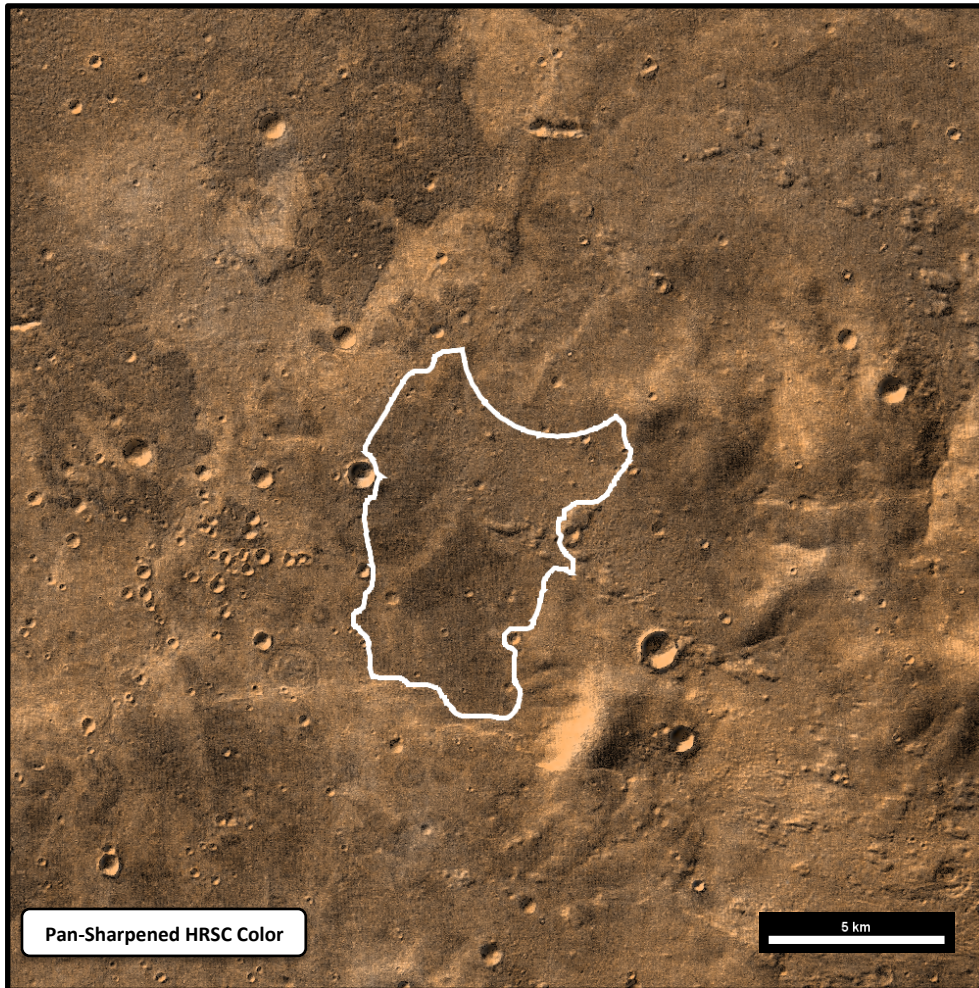
**Coverage:** THEMIS, HRSC, CTX

- Landing site of  $\sim 25\text{km}^2$
- Meets all landing site requirements
- Intersects SROI-5 and RROI-5



# Habitation Site

1<sup>st</sup> EZ Workshop for Human Missions to Mars



**Location:** -37.427N, 350.590E

**Elevation:** 1,189m – 1,246m

**Coverage:** THEMIS, HRSC, CTX

- Habitation site of  $\sim 50\text{km}^2$
- Meets all habitation site requirements
- Adjacent to landing site